

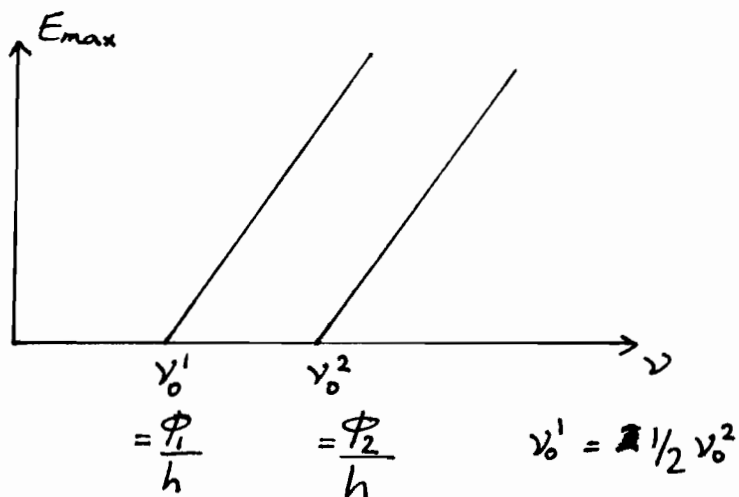
Physics IV - Mock Exam

SS 2007 - SOLUTIONS.

04/07/2007

①

a)



$$E_{\max} = h(\nu - \nu_0)$$

main features:

- the two lines have the same slope  $h$
- the intercept with the x-axis gives work function  $\Phi = \nu_{\text{intercept}} \times h$
- if incident photons do not have energy  $> \Phi$ , they cannot excite electrons out of the metal  $\Rightarrow$  zero current in photoelectric experiment &  $E_{\max} = 0$
- maximum energy that the electrons can have is  $h\nu - \Phi$  since one photon is absorbed each electron.

②  $\Phi_{\text{tungsten}} = 5.4 \text{ eV}$      $\lambda = 175 \text{ nm}$      $E_{\max} = 1.7 \text{ eV}$

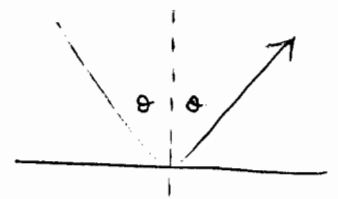
$$E_{\max} = h\nu - \Phi \quad \Rightarrow \quad h = \frac{E_{\max} + \Phi}{\nu}$$

$$\begin{aligned} \nu &= \frac{c}{\lambda} \quad \Rightarrow \quad h = \frac{(E_{\max} + \Phi) \lambda}{c} = \frac{(1.7 + 5.4) \times 175 \times 10^{-9}}{3 \times 10^8} \text{ eV}\cdot\text{s} \\ &= 4.14 \times 10^{-15} \text{ eV}\cdot\text{s} \\ &= 6.63 \times 10^{-34} \text{ J}\cdot\text{s} \end{aligned}$$

②  $E_{\text{electron}} = 50 \text{ eV}$       $a = 0.22 \text{ nm}$

Bragg reflection      $2a \cos \theta = n\lambda$

Need to know  $\lambda_{\text{electron}}$  :-



[NB:  $\cos \theta$  when  $\theta$  is measured from the normal.]

de Broglie  $p = h/\lambda$      and  $E = p^2/2m$

$$\Rightarrow E = \frac{h^2}{2m\lambda^2} \quad \Rightarrow \lambda_{\text{electron}} = \sqrt{\frac{h^2}{2mE_{\text{electron}}}}$$

$$\theta = \cos^{-1} \left( \frac{\lambda}{2a} \right) = \cos^{-1} \left[ \frac{h}{2a} \frac{1}{\sqrt{2mE_{\text{electron}}}} \right]$$

(for first maximum)  $n=1$       $= \underline{\underline{67^\circ}}$

⑥ X-rays must have the same wavelength as the electrons.

$$\lambda = \frac{h}{\sqrt{2mE_{\text{electron}}}} = 1.74 \times 10^{-10} \text{ m}$$

energy of X rays (photons)  $E = \frac{hc}{\lambda} = \frac{6.63 \times 10^{-34} \times 3 \times 10^8}{1.74 \times 10^{-10}}$

$$= 1.145 \times 10^{-15} \text{ J}$$

$$= 7.2 \text{ keV}$$

③ The combined wavefunction of two fermions must be antisymmetric with respect to changing the positions of the two fermions.

ie. Swapping the positions of the two fermions inverts the sign of the wavefunction.

Proof that two fermions cannot occupy the same state :-

- representing single particle state  $i$  by  $\psi_i$
- labelling different fermions as A, B.  $\Rightarrow \psi_i^A$  means fermion A in state  $i$ .

Total wavefunction must be antisymmetric :-

$$\psi_{\text{total}} = \frac{1}{\sqrt{2}} (\psi_i^A \psi_j^B - \psi_j^B \psi_i^A)$$

When  $i = j$ ,

$$\psi_{\text{total}} = \frac{1}{\sqrt{2}} (\psi_i^A \psi_i^B - \psi_i^B \psi_i^A) = 0$$

$\Rightarrow$  two fermions cannot exist simultaneously in the same single particle state.

④ 3D box with as hard walls.

①

$$\Psi = A \sin\left(\frac{n_x \pi x}{a}\right) \sin\left(\frac{n_y \pi y}{a}\right) \sin\left(\frac{n_z \pi z}{a}\right)$$

$$\int_0^a \int_0^a \int_0^a \Psi^* \Psi \, dx \, dy \, dz = 1$$

$$\Rightarrow A^2 \int_0^a \sin^2\left(\frac{n_x \pi x}{a}\right) dx \int_0^a \sin^2\left(\frac{n_y \pi y}{a}\right) dy \int_0^a \sin^2\left(\frac{n_z \pi z}{a}\right) dz = 1$$

$$\int_0^a \sin^2\left(\frac{n \pi x}{a}\right) dx = \int_0^a \frac{1}{2} \left(1 - \cos \frac{2n \pi x}{a}\right) dx = \left[ \frac{x}{2} - \frac{1}{2} \left(\frac{a}{2n \pi}\right) \sin\left(\frac{2n \pi x}{a}\right) \right]_0^a$$

$$= \frac{a}{2} - \frac{1}{2} \left(\frac{a}{2n \pi}\right) \sin(2n \pi) = \frac{a}{2}$$

↳  $n$  is an integer  $\Rightarrow \sin(2n \pi) = 0$

$$\int_0^a \int_0^a \int_0^a \Psi^* \Psi \, dx \, dy \, dz = A^2 \left(\frac{a}{2}\right)^3 = 1$$

$$\Rightarrow A = \frac{2\sqrt{2}}{a^{3/2}}$$

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④

⑥ time-indep. Schrödinger equation:  $H\psi = E\psi$

$$-\frac{\hbar^2}{2m} \nabla^2 \psi + V\psi = E\psi$$

Inside the box  $V=0 \Rightarrow -\frac{\hbar^2}{2m} \nabla^2 \psi = E\psi$

$$\begin{aligned} \nabla^2 \psi &= \left( \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) \psi = \left( -\frac{n_x^2 \pi^2}{a^2} - \frac{n_y^2 \pi^2}{a^2} - \frac{n_z^2 \pi^2}{a^2} \right) \psi \\ &= -\frac{\pi^2}{a^2} (n_x^2 + n_y^2 + n_z^2) \psi \end{aligned}$$

$$\Rightarrow E = \frac{\hbar^2 \pi^2}{2ma^2} (n_x^2 + n_y^2 + n_z^2)$$

⑦ possible values for  $n_x, n_y, n_z$ :

	$n_x$	$n_y$	$n_z$	$E$
ground state ←	1	1	1	$\frac{3\hbar^2 \pi^2}{2ma^2}$
3 different wavefunctions x 2 for spin ← 1st excited state	2	1	1	$\frac{3\hbar^2 \pi^2}{ma^2}$
	1	2	1	$\frac{3\hbar^2 \pi^2}{ma^2}$
	1	1	2	$\frac{3\hbar^2 \pi^2}{ma^2}$
	2	2	1	$\frac{9\hbar^2 \pi^2}{2ma^2}$
		⋮		⋮

⇒ degeneracy of the 1st excited state is 6

④

①

Particle is in the ground state:  $n_x = n_y = n_z = 1$

$$\text{Probability } P = \int_0^{a/4} \int_0^{a/4} \int_0^{a/4} \psi^* \psi \, dx \, dy \, dz$$

$$\int_0^{a/4} \sin^2\left(\frac{\pi x}{a}\right) dx = \int_0^{a/4} \frac{1}{2} \left(1 - \cos\left(\frac{2\pi x}{a}\right)\right) dx$$

$$= \left[ \frac{x}{2} - \frac{1}{2} \left(\frac{a}{2\pi}\right) \sin\left(\frac{2\pi x}{a}\right) \right]_0^{a/4}$$

$$= \frac{a}{8} - \frac{a}{4\pi} \sin\left(\frac{\pi}{2}\right) = a \left(\frac{1}{8} - \frac{1}{4\pi}\right)$$

$$\Rightarrow P = \frac{8}{a^3} \times \left(\frac{a}{8} - \frac{a}{4\pi}\right)^3 = \frac{\left(\frac{1}{4} - \frac{1}{2\pi}\right)^3}{A^2} = \underline{\underline{0.00075}}$$

Classically the probability of finding the particle in a volume  $V$  is the ratio of  $V$  to the total volume, therefore

$$P_{\text{classical}} = \frac{(a/4)^3}{a^3} = \left(\frac{1}{4}\right)^3 = 0.015$$

$\Rightarrow$  the QM probability is much lower due to the sin shape of the wave-function.

$$\textcircled{5} \quad i\hbar \frac{d}{dt} \psi = H \psi$$

$$\textcircled{a} \quad \psi(t=0) = \psi_1$$

$$\psi(t) = a(t) \psi_1 + b(t) \psi_2$$

$$i\hbar \left( a'(t) \psi_1 + b'(t) \psi_2 \right) = a(t) E_1 \psi_1 + b(t) E_2 \psi_2$$

$\psi_1$  &  $\psi_2$  orthogonal  $\therefore$  separate variables :-

$$\left\{ \begin{array}{l} i\hbar a'(t) = E_1 a(t) \\ i\hbar b'(t) = E_2 b(t) \end{array} \right\} \quad \begin{array}{l} a(0) = 1 \\ b(0) = 0 \end{array}$$

$$\int_{a(0)}^{a(t)} \frac{1}{a} da = \int_0^t \frac{-iE_1}{\hbar} dt \quad \& \text{ similar for } b(t)$$

$$\Rightarrow a(t) = a(0) e^{-iE_1 t/\hbar} = e^{-iE_1 t/\hbar}$$

$$b(t) = 0$$

$$\Rightarrow \psi(t) = e^{-iE_1 t/\hbar} \psi_1 \quad \text{//}$$

$$\textcircled{b} \quad A \text{ measured at } t=0 \quad \phi_1 = \frac{1}{\sqrt{2}} (\psi_1 + i\psi_2)$$

$$\phi_2 = \frac{1}{\sqrt{2}} (\psi_1 - i\psi_2)$$

$$\Rightarrow \psi_1 = \frac{1}{\sqrt{2}} (\phi_1 + \phi_2) = x \phi_1 + y \phi_2$$

$$\psi_2 = \frac{1}{\sqrt{2}} (\phi_2 - \phi_1)$$

Probabilities of measuring  $a_1$ ,  $P(a_1) = |x|^2 = \frac{1}{2}$  //

" " "  $a_2$ ,  $P(a_2) = |y|^2 = \frac{1}{2}$  //



③  $a_1$  is the measurement result  $\Rightarrow$

$$\psi(0) = \phi_1 = \frac{1}{\sqrt{2}}(\psi_1 + i\psi_2)$$

$$\psi(t) = \frac{1}{\sqrt{2}}(e^{-iE_1 t/\hbar} \psi_1 + i e^{-iE_2 t/\hbar} \psi_2)$$

$$= \frac{1}{2} (e^{-iE_1 t/\hbar} (\phi_1 + \phi_2) + i e^{-iE_2 t/\hbar} (\phi_2 - \phi_1))$$

$$= \frac{1}{2} ((\phi_1 + \phi_2) + (\phi_1 - \phi_2) e^{-i(E_2 - E_1)t/\hbar}) e^{-iE_1 t/\hbar}$$

↓  
ignore  
"global"  
phase

$$\psi(t) = \frac{1}{2} ((1 + e^{-i\omega t}) \phi_1 + (1 - e^{-i\omega t}) \phi_2)$$

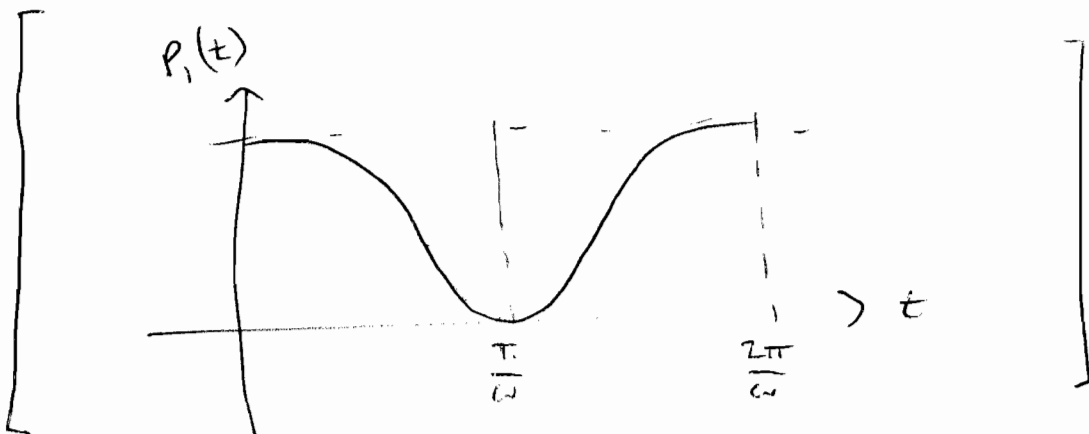
$$= x \phi_1 + y \phi_2 \quad \text{where } \omega = (E_2 - E_1)/\hbar$$

$$P_1(t) = |x|^2$$

$$= \frac{1}{4} (1 + e^{-i\omega t})(1 + e^{i\omega t})$$

$$= \frac{1}{4} (1 + 1 + e^{-i\omega t} + e^{i\omega t})$$

$$= \frac{1}{2} (1 + \cos \omega t) //$$



⑥ Zeeman Effect  $\Delta E = \frac{eB}{2m_e} (L_z + 2S_z)$

①

$n=1$  energy levels : 1s only.

$n=2$  energy levels : 2s and 2p.

$$L_z = \hbar m_l$$

$$S_z = \hbar m_s$$

1s :-  $m_l = 0, m_s = \pm 1/2 \Rightarrow \Delta E_{1s} = \pm \frac{eB\hbar}{2m_e}$

2s :-  $m_l = 0, m_s = \pm 1/2 \Rightarrow \Delta E_{2s} = \pm \frac{eB\hbar}{2m_e}$

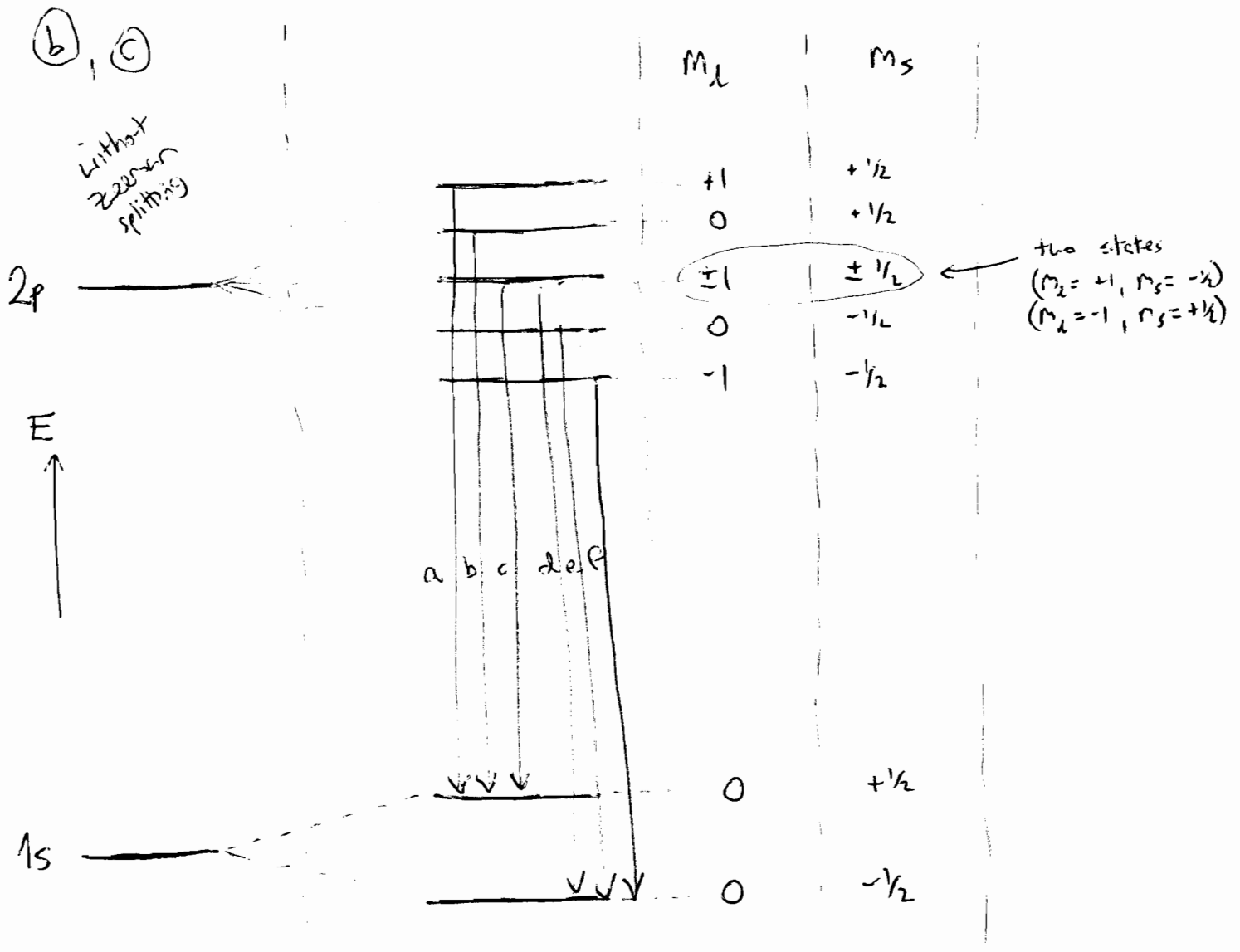
2p :-  $m_l = 0, \pm 1, m_s = \pm 1/2$

$$\Rightarrow \Delta E_{2p} = \frac{eB\hbar}{2m_e} \times \begin{matrix} m_l & m_s \\ \left. \begin{matrix} -1 \\ -1 \\ 0 \\ 0 \\ +1 \\ +1 \end{matrix} \right\} + 2 \times \left. \begin{matrix} +1/2 \\ -1/2 \\ +1/2 \\ -1/2 \\ +1/2 \\ -1/2 \end{matrix} \right\} = \frac{eB\hbar}{2m_e} \times \begin{matrix} 0 \\ -2 \\ +1 \\ -1 \\ +2 \\ 0 \end{matrix}$$

$$= (-2, -1, 0, +1, +2) \times \frac{eB\hbar}{2m_e} \quad \left( \begin{array}{l} 6 \text{ different states,} \\ \text{but 5 different} \\ \text{energies.} \end{array} \right)$$

↑ degeneracy 2

pts for ① and ②



Allowed transitions :-  $m_s$  conserved.  
 6 transitions shown above, labelled a-f.

Energy of transitions :  $\Delta E_a = \Delta E_d$   
 $\Delta E_b = \Delta E_e$   
 $\Delta E_c = \Delta E_f$

$\Rightarrow$  (3) spectral lines only.

(d)  $B = 10T$

Splitting is  $\Delta E = \frac{eB\hbar}{2m_e} = h\nu$

$\nu = \frac{eB}{4\pi m_e} = 140 \text{ GHz}$

$$\textcircled{7} \quad E_J = \frac{L^2}{2I} = \frac{J(J+1) \hbar^2}{2\mu r^2}$$

$$\mu = \frac{m_1 m_2}{m_1 + m_2} \quad \text{where } m_1 = m_u \text{ (Hydrogen)}$$

$$m_2 = 19 m_u \text{ (Fluorine)}$$

$$\Rightarrow \mu = \frac{19}{20} m_u$$

lowest frequency spectral line  $\Rightarrow$  lowest energy transition

$\rightarrow J=0$  to  $J=1$  transition

$$\Delta E = \frac{1(1+1) \hbar^2}{2 \cdot \frac{19}{20} m_u r^2}$$

& need to find  $r$ .

$$r^2 = \frac{\hbar^2}{\frac{19}{20} m_u \Delta E}$$

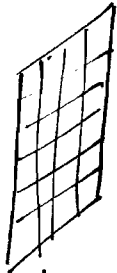
$$\& \Delta E = h\nu$$

where  $\nu = 1.25 \times 10^{12} \text{ Hz}$ .

$$\Rightarrow r = \sqrt{\frac{\hbar^2}{\frac{19}{20} m_u h\nu}} = \sqrt{\frac{\hbar}{\frac{19}{20} m_u \cdot 2\pi \cdot \nu}}$$

$$= 9.2 \times 10^{-11} \text{ m} //$$

8



Solar panel  
on Earth

$$\text{area } A = 10 \text{ m}^2$$

$$\text{efficiency} = 30\%$$

$$\text{Earth to Sun distance } d_{E-S} = 150 \times 10^6 \text{ km}$$

$$\text{Sun's temperature } T_S = 6000 \text{ K}$$

$$\text{Sun's radius } R_S = 7 \times 10^5 \text{ km}$$

total power per unit area emitted from a black-body at temperature  $T$  is given by the Stefan-Boltzmann law:

$$\frac{P}{\text{unit area}} = \sigma T^4$$

treating the Sun as a black-body, the total power emitted by the Sun is:

$$P_{\text{total}} = \sigma T_S^4 \times 4\pi R_S^2$$

$\Rightarrow$  power per unit area falling on Earth from the Sun is:

$$\frac{P_{\text{on Earth}}}{\text{unit area}} = \frac{P_{\text{total}}}{4\pi d_{E-S}^2} = \sigma T_S^4 \times \frac{R_S^2}{d_{E-S}^2}$$

the total power output of the Solar Panel is then given by:

$$\begin{aligned}
 & \underset{\substack{\swarrow \\ \text{efficiency}}}{0.3} \times \underset{\substack{\downarrow \\ \text{area of} \\ \text{the panel}}}{A} \times \sigma T_S^4 \times \frac{R_S^2}{d_{E-S}^2} = 0.3 \times 10 \times 5.67 \times 10^{-8} \times 6000^4 \times \left( \frac{7 \times 10^5}{150 \times 10^6} \right)^2 \\
 & = \underline{\underline{4.8 \text{ KW}}}
 \end{aligned}$$